

Astronomical Observation

天文觀測



- To study the properties of celestial objects by means of observations

以觀測的手段瞭解宇宙天體的性質，探討其來源、演化，以及彼此之交互作用

Star → EM radiation → Detection → Analysis

Properties of stars ← Interpretation

宇（上下四方） 宙（古往今來）

空間

時間

天球 (celestial sphere)

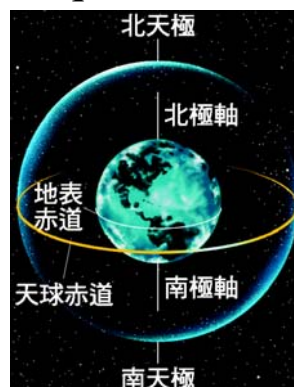
--- 角度 (angle)

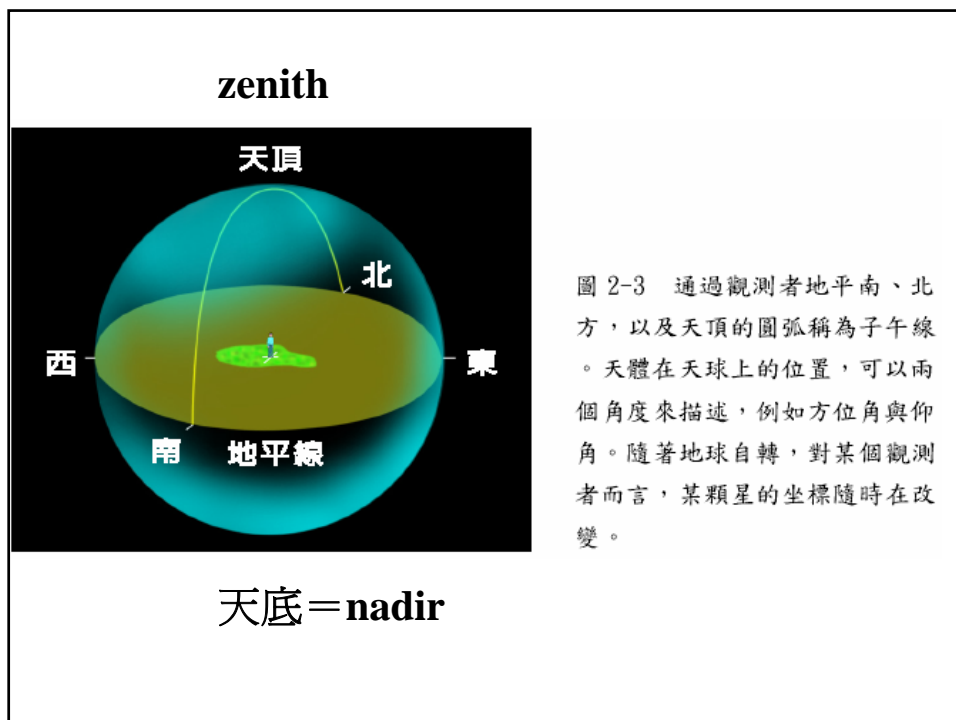
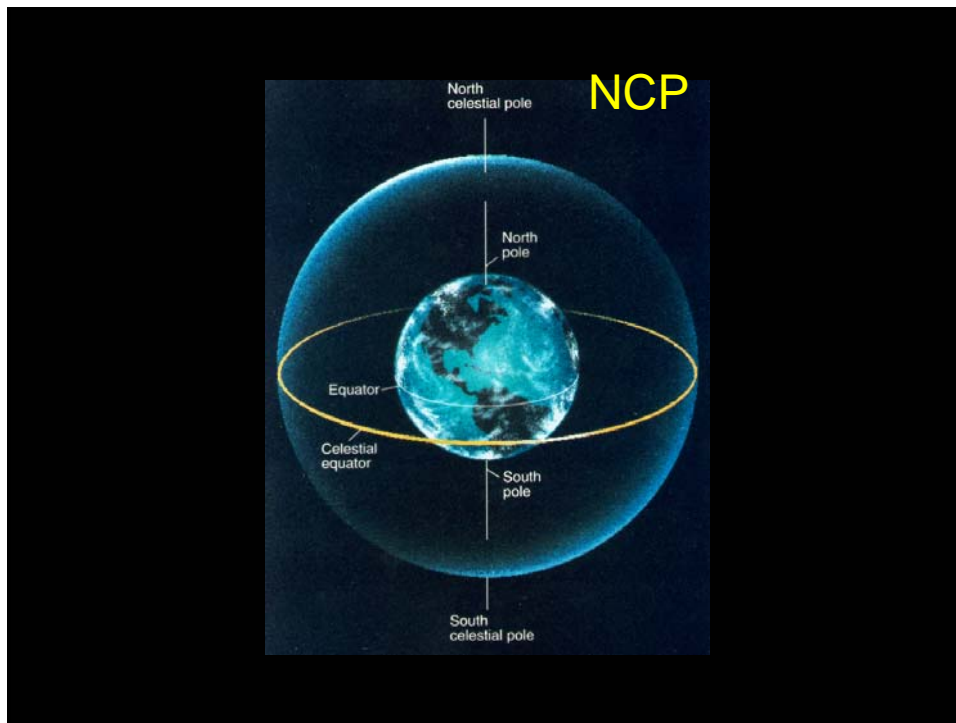
距離（深度、沿視線方向；
徑向）

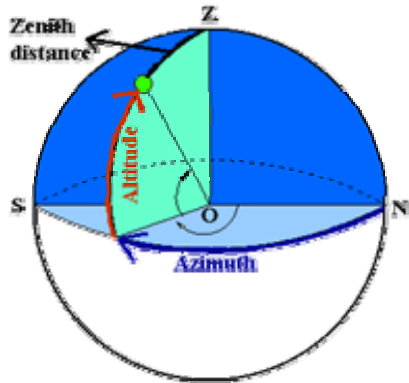
Action, reaction, interaction

因果關係

(con)sequences







altitude (仰角)

(= elevation)

-90° to $+90^\circ$

azimuth (方位角)

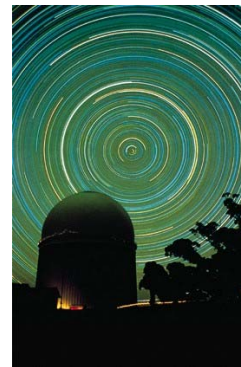
e.g., measured from north,
clockwise

0° to 360°

Zenith distance = $90^\circ - \text{Alt}$

Measuring the Sky

- Unit of distance in astronomy
 - 1 astronomical unit (AU) 天文單位
 $= 149.6 \times 10^6 \text{ km}$
 約一億五千萬公里
 - 1 light year (ly) $= 9.460 \times 10^{12} \text{ km}$
 $\sim 6 \times 10^4 \text{ AU}$
 - 1 parsec (pc) = 206,265 AU 秒差距
 $= 3.086 \times 10^{13} \text{ km} \sim 3.26 \text{ ly}$
- Unit of separation (on the sky)
 - 1 degree = 60 minutes of arc (arcminutes) = $60'$
 - 1 arcminute = 60 arcseconds = $60''$



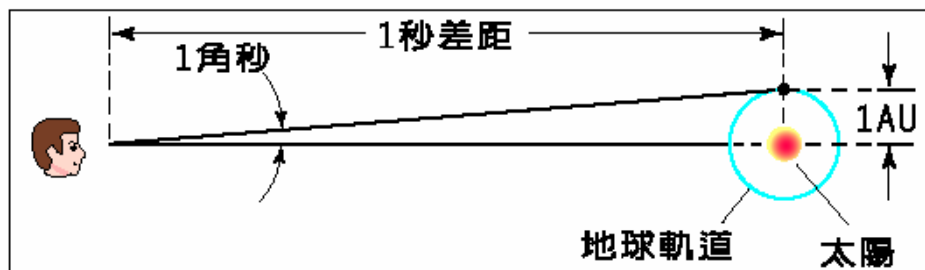
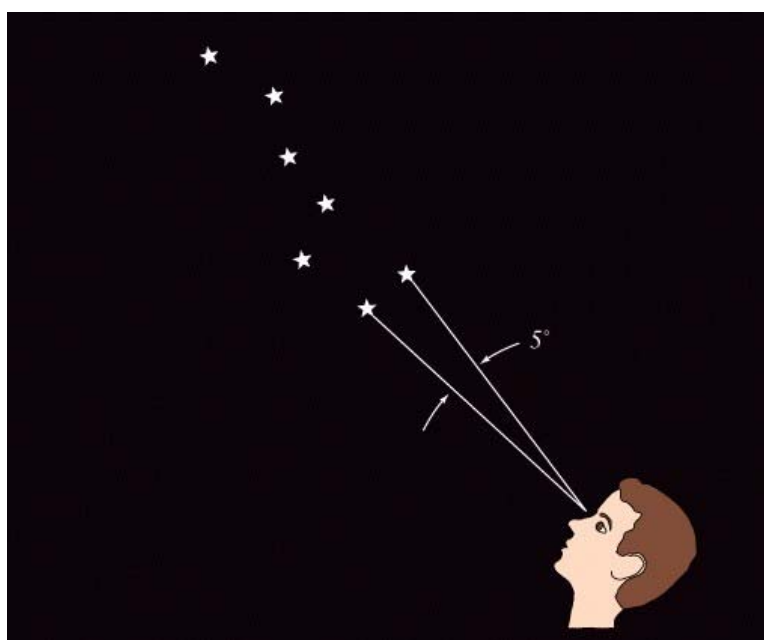
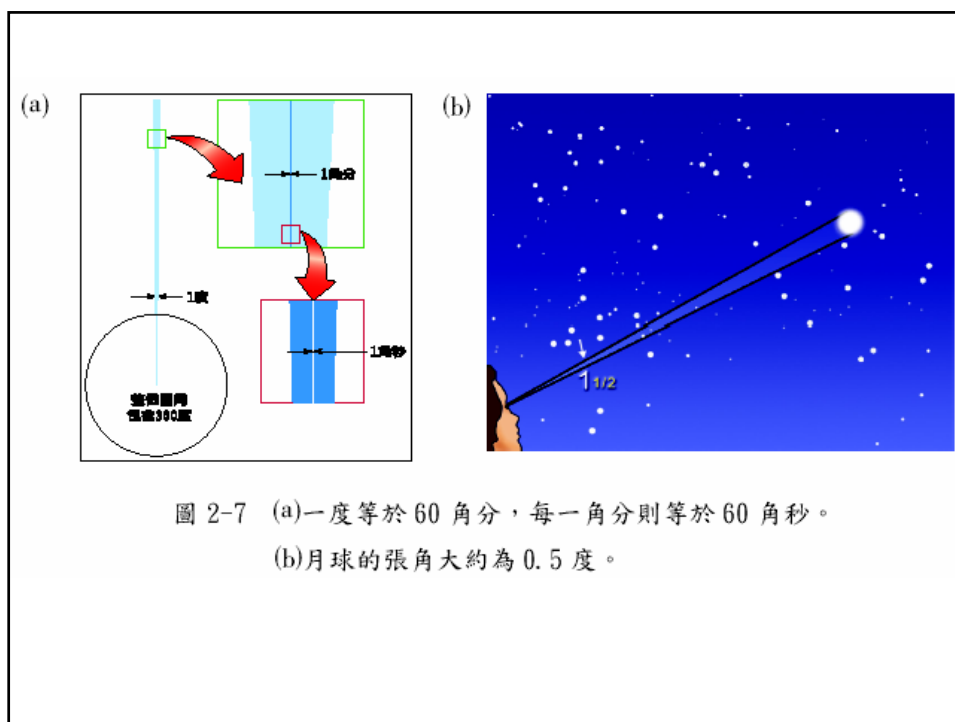
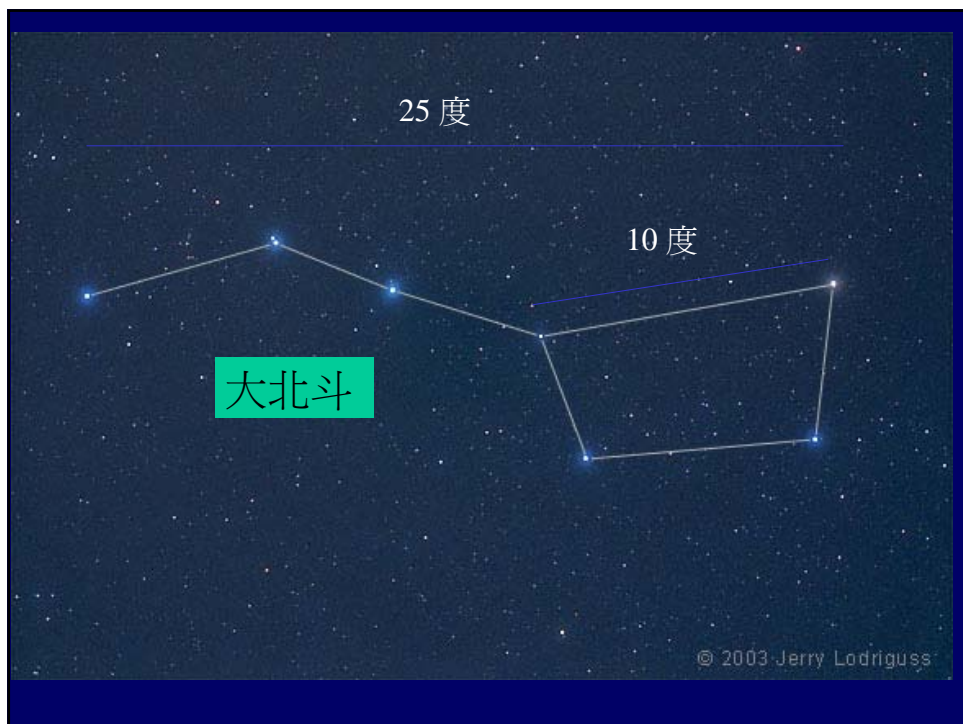
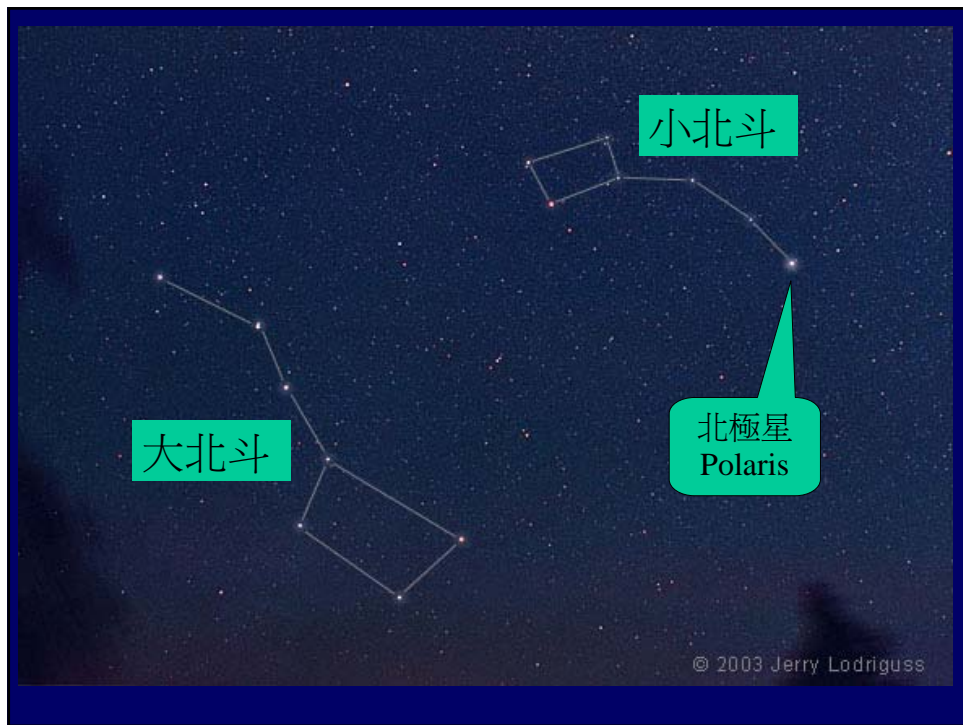


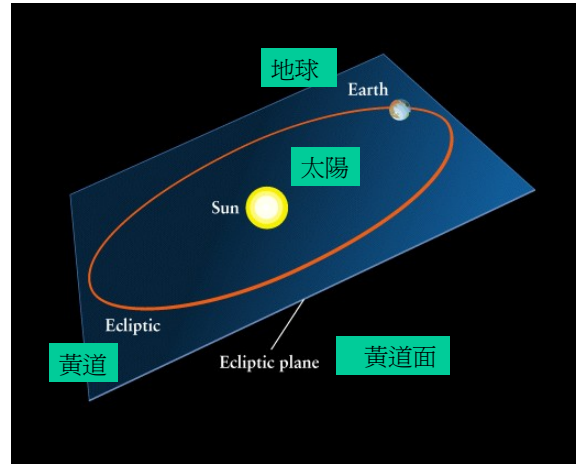
圖 2-9 當相離 1 pc 時，1 AU 的張角為 1 角秒。







Earth's Revolution 地球公轉



The Earth revolves around the Sun. The passage is called the **ecliptic** (黃道).

黃道面：The ecliptic plane intercepts the celestial equator with an angle of 23.4 degrees.

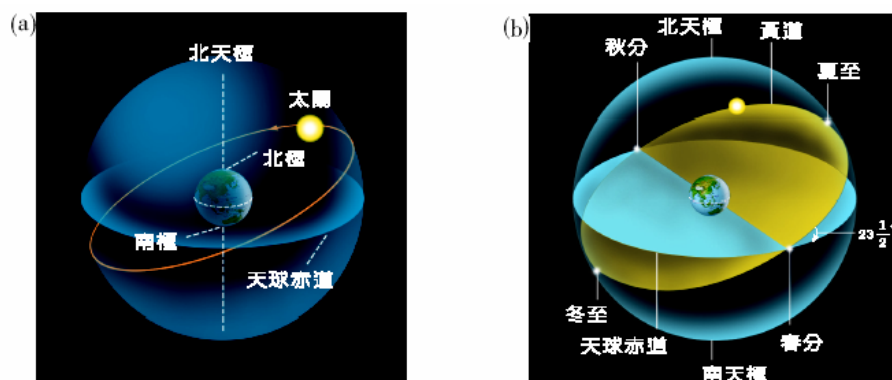


圖 2-12 (a)從地球看太陽一年當中的運動，最北時達到赤緯北方 23.5 度，然後通過天球赤道，繼續南行直到赤緯南方 23.5 度。

(b)黃道面與天球赤道成 23.5 度傾角，兩個面的交點分別稱為春分點及秋分點，黃道面最（上）北方的點稱為夏至，最南方的點則稱為冬至。

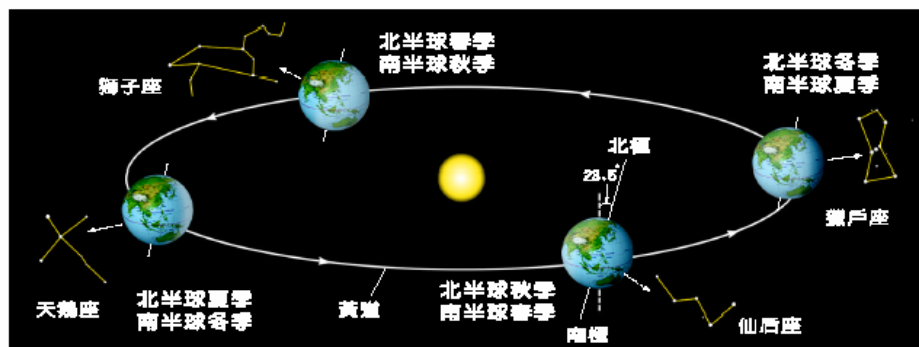
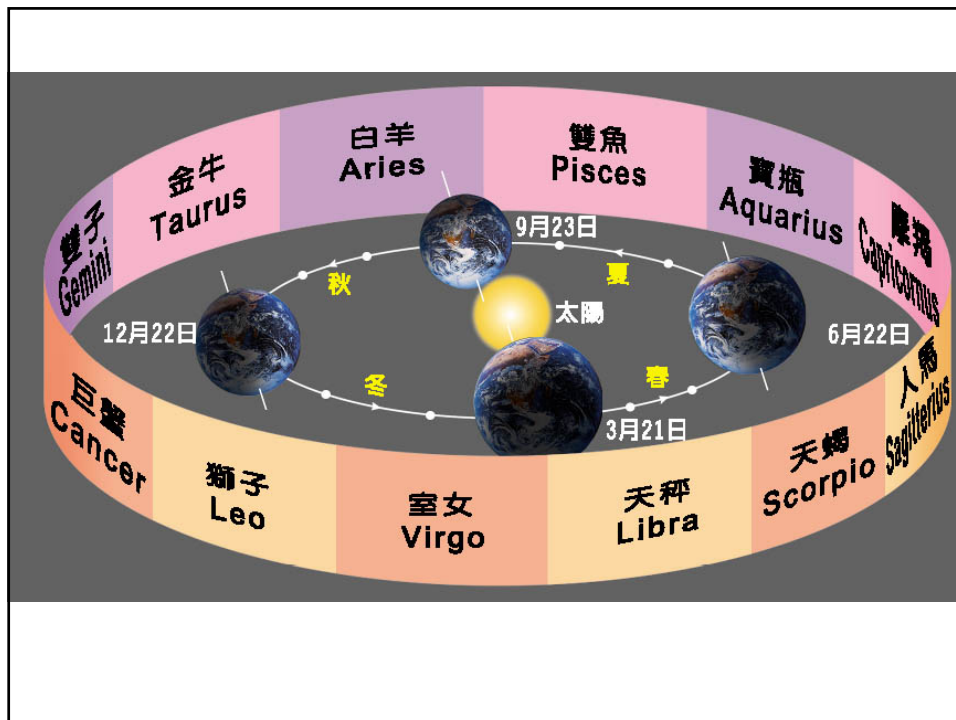


圖 2-11 黃道面與天球赤道成 23.5 度傾角，兩個面的交點分別稱為春分點及秋分點，黃道面最（上）北方的點稱為夏至，最南方的點則稱為冬至。

At different hours in a night, one sees different night sky. At different months one sees also different sky → 黃道十三宮 (zodiac signs)

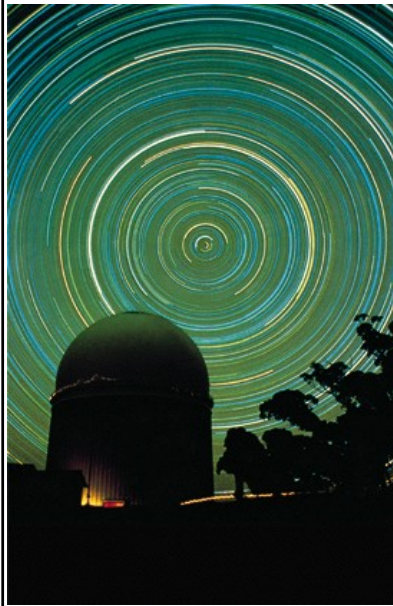
Table 1-1
The 13 Constellations of the Zodiac

Constellation	Dates of Sun's Passage Through
Pisces	March 13–April 20
Aries	April 20–May 13
Taurus	May 13–June 21
Gemini	June 21–July 20
Cancer	July 20–August 11
Leo	August 11–September 18
Virgo	September 18–November 1
Libra	November 1–November 22
Scorpius	November 22–December 1
missing → Ophiuchus 蛇夫座	December 1–December 19
Sagittarius	December 19–January 19
Capricorn	January 19–February 18
Aquarius	February 18–March 13

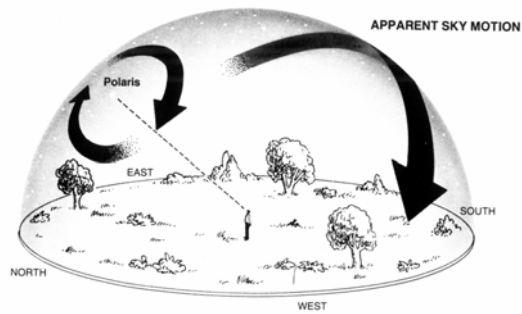


Earth's Rotation (Spin) 地球自轉

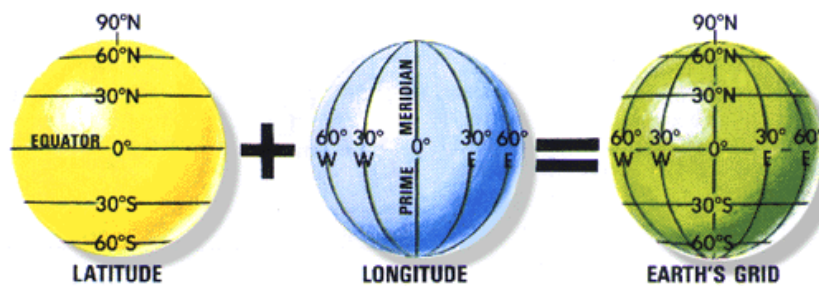
- Counterclockwise (west → east) if viewing from the above (the north pole) → Sun, moon and stars rise in the east and set in the west
- Earth's orbital motion also counterclockwise
- Extending Earth's spin axis
→ 天 (球) 北極 (**north celestial pole**; NCP)
NCP appears not to spin; all other stars seem to move around the NCP or SCP in a day
- 北極星 (**Polaris**) : closest to NCP now ($< 1^\circ$)
- Babylonians 3000 years ago 360 days
→ 360 deg in a circle



In the duration of a day, all stars seem to move around the NCP. The stars actually do not move. The Earth's spin causes the apparent motion of stars.



The coordinate System on the Surface of the Earth — Longitude and Latitude



- **Longitude and Latitude**

Greenwich longitude = 0
格林威治當地時間
(local time)

→ Universal Time
(UT ; 國際標準時)

Taiwan Time = UT + 8 hr

赤道 latitude = 0
北極：北緯90度；
南極：90°S or -90°N

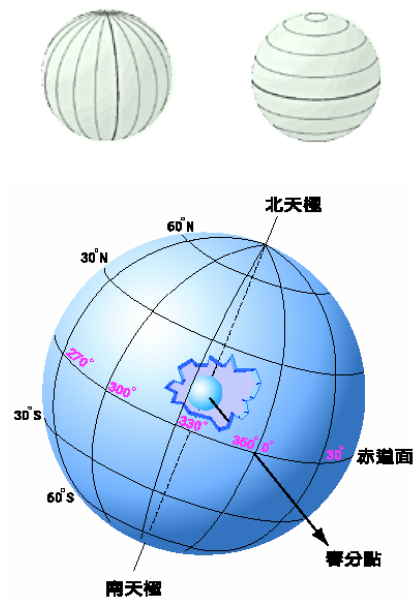
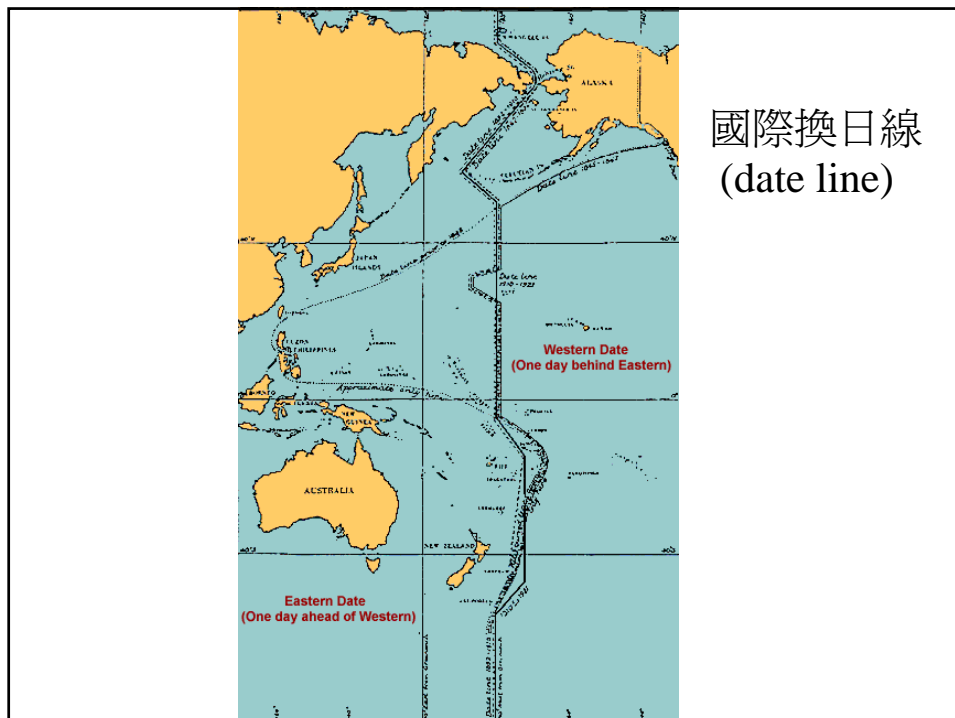
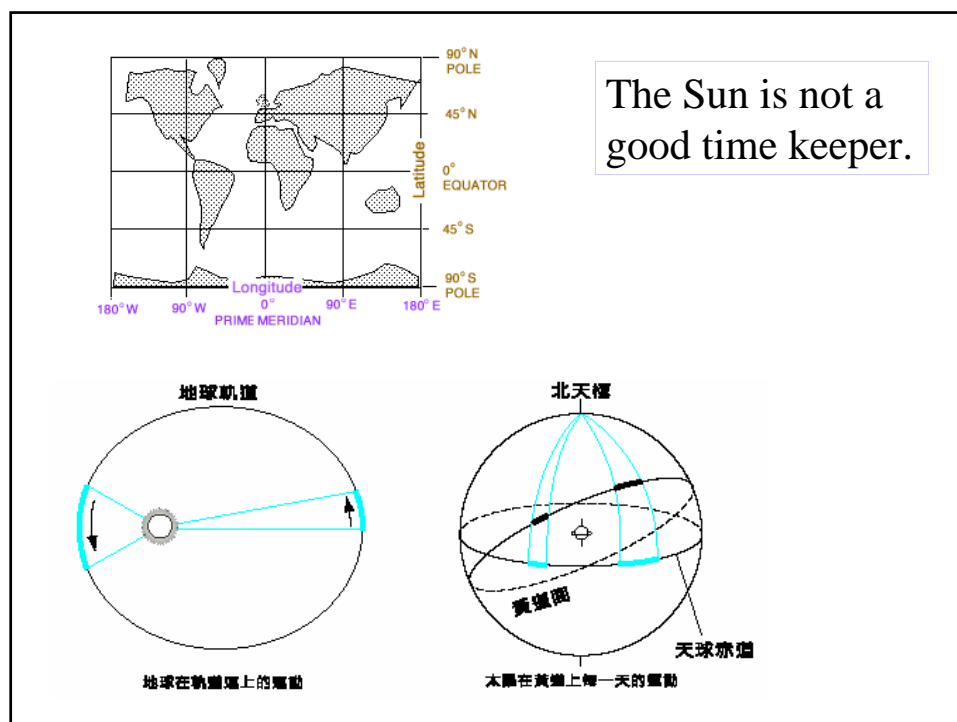
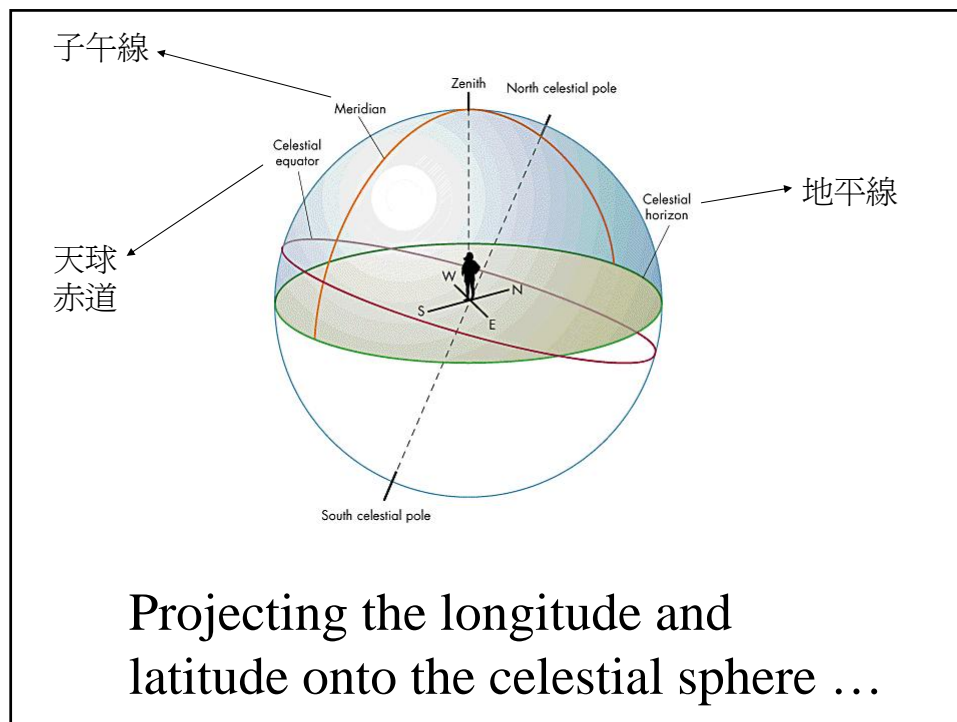


圖 2-6 把地表的經度與緯度向外擴張，便能將類似的坐標系統應用在天球上，用以描述天體的位置



國際換日線
(date line)



- **Sidereal Time** (恆星時)

Time as measured by reference to the stars
= hour angle of the **vernal equinox**
= right ascension of stars currently on the observer's **meridian**

- **Solar Time** (太陽時)

Time with respect to the sun
= hour angle of the sun plus 12 hours

apparent solar time → true sun, irregularly

mean solar time → for timekeeping

Solar time loses about 4 minutes a day against sidereal time; a star rises **4 minutes earlier** a day

Q : Which is longer, a solar day or a sidereal day? By how much?

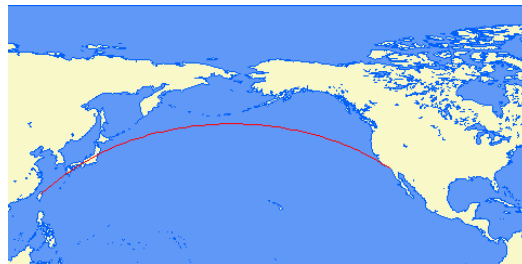
1. Solar day; one hour
2. Solar day; 4 minutes
3. Sidereal day; one hour
4. Sidereal day; 4 minutes

- The sidereal day is 23 h 56 m 04.1 s
- Each apparent solar day in a year is different, with a variation of ~30 s
- One mean solar day, by definition, is exactly 24 hours.
- The mean sidereal day is about 4 minutes shorter than the mean solar day.
→ 相同的夜空隔天會早約四分鐘升起
- Other sidereal periods
sidereal month for the Moon = 27.32166 d
c.f. (mean) synodic month 朔望月 = 29.53059 d
sidereal day; sidereal year

大圓 (great circle)

從台北到洛杉磯
的飛行路線？

子午線 (meridian)



赤經 (right ascension; RA)

通常以時間單位表示

天球赤道與黃道

交點之一 RA=0h

每年約3月21日

太陽行至春分點

赤緯 (declination; DEC)

通常以角度單位表示



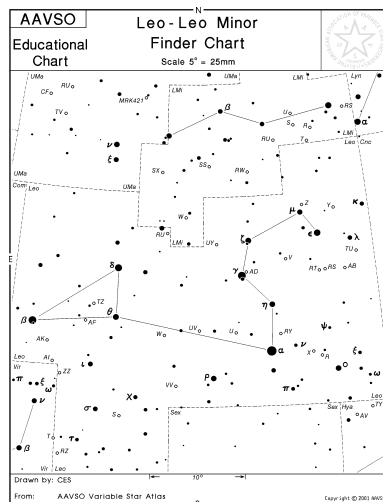
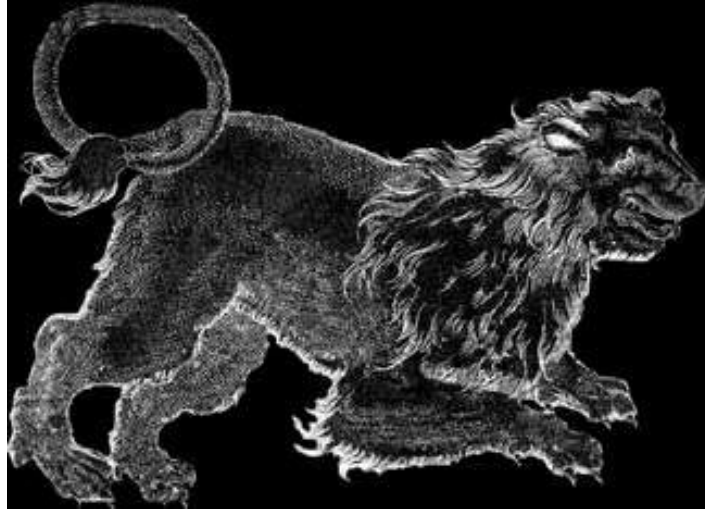
TPE-LAX

TPE-NYC

星座 (constellation) --- Stars in a constellation may not be associated with each other. They are at different distances but merely projected onto the same part of the sky. There are a total of 88 constellations for the whole sky.



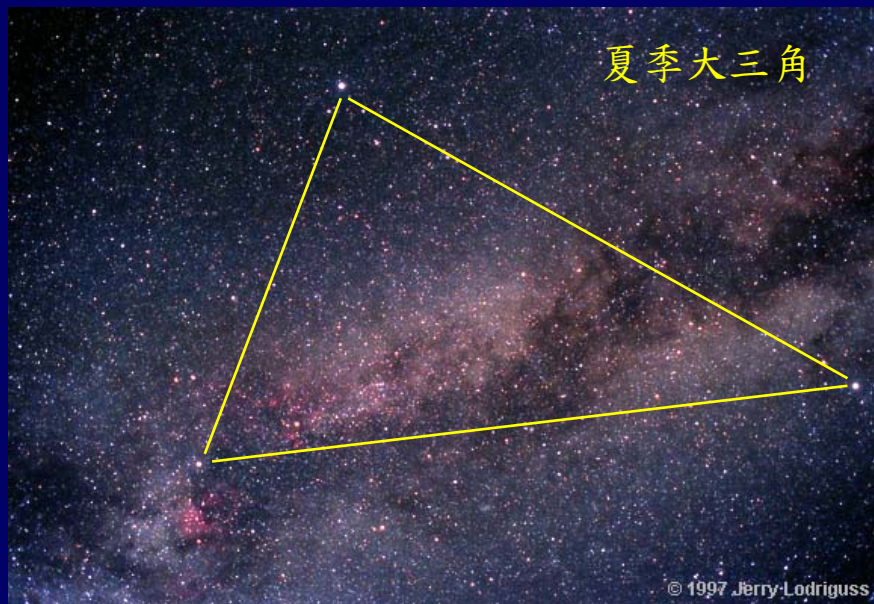
The constellation Leo “Lion” appears in the spring sky.

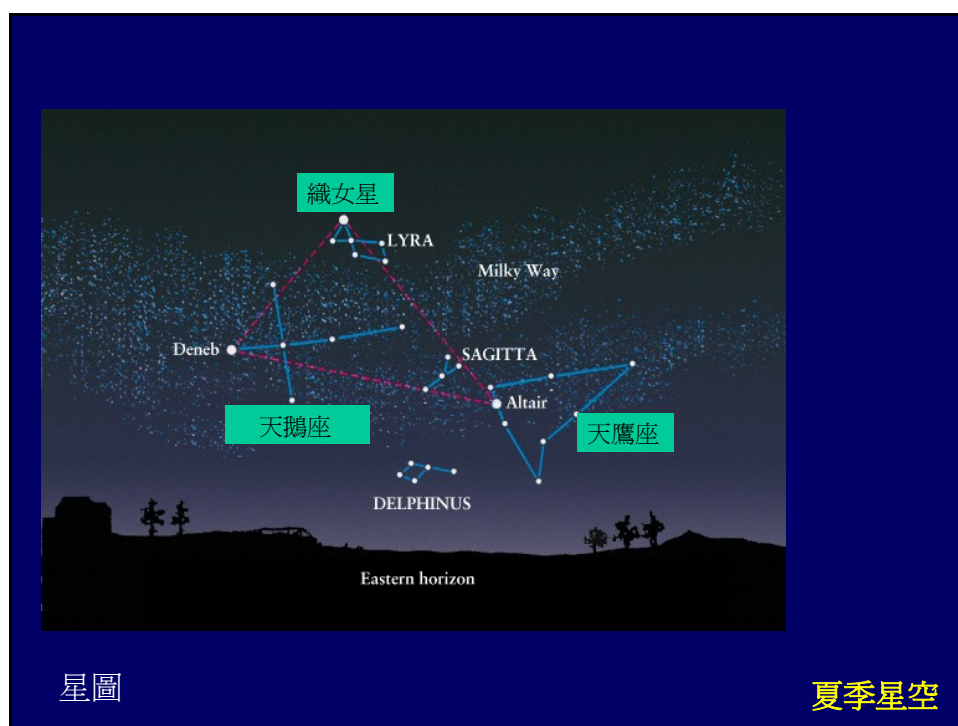


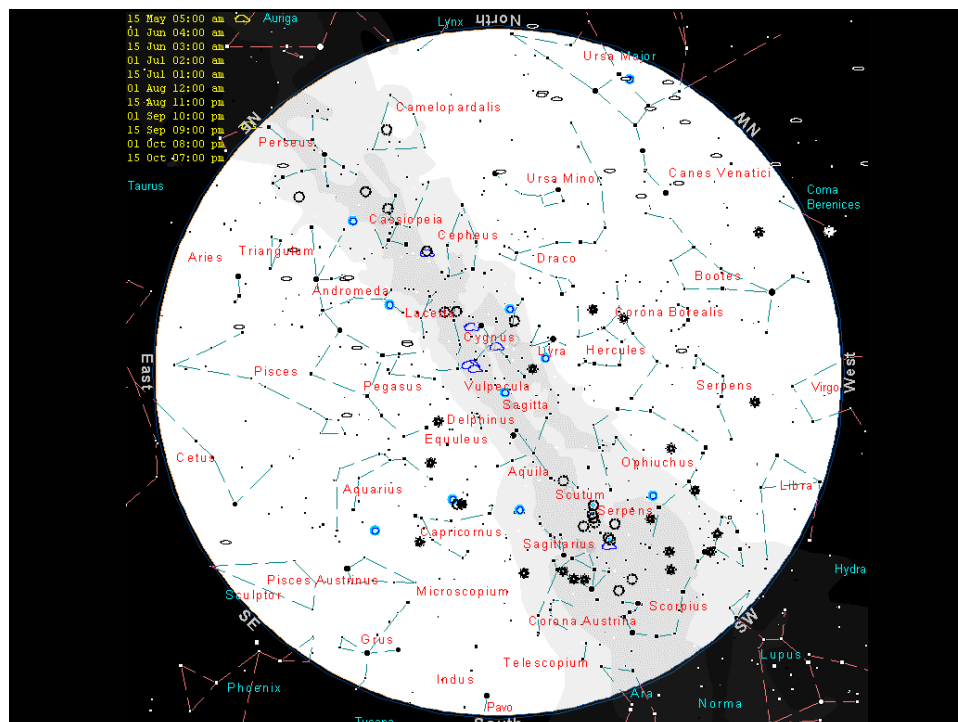
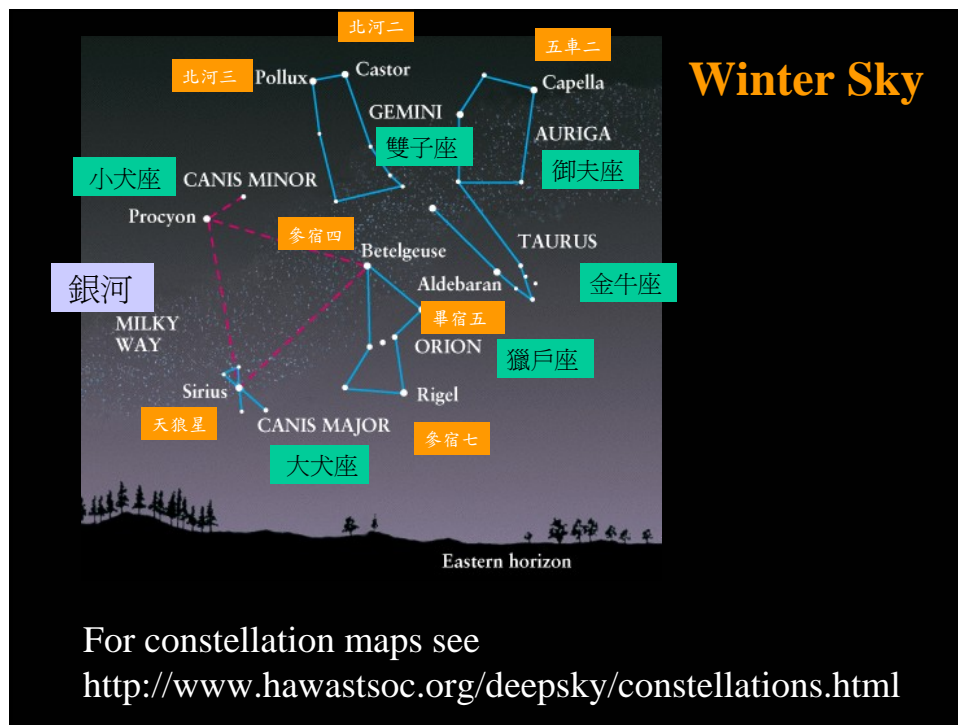
A constellation is more than a “star pattern”. It stands for a patch (region) of the sky. A star chart (星圖) shows Leo and its neighboring constellations.

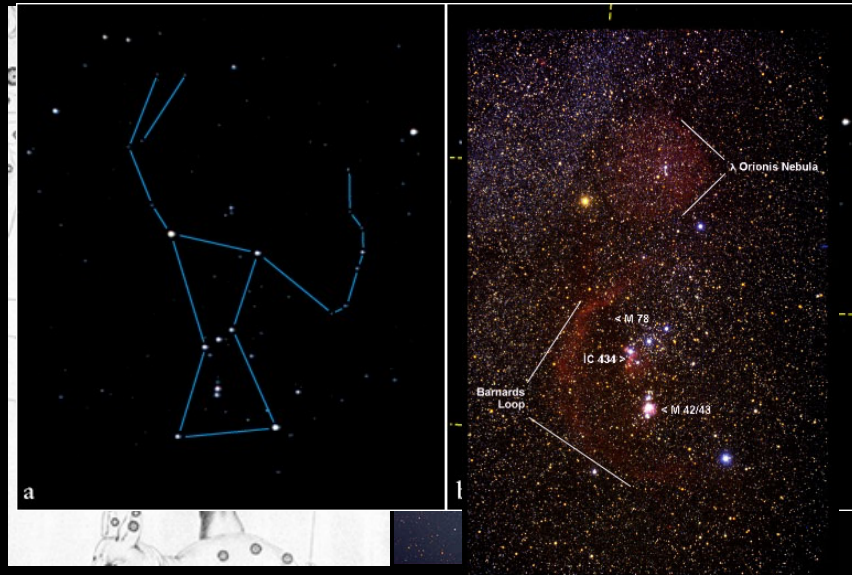


Constellation Cygnus (photograph)







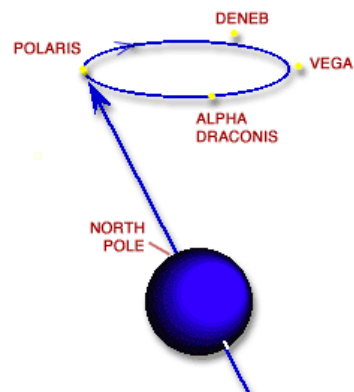


Orion 獵戶座

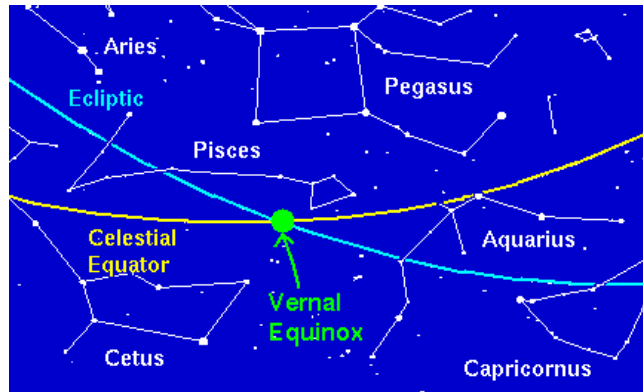
<http://www.allthesky.com/constellations/orion/mainm.html>

Celestial Coordinates

- The sky position of a star is specified by its RA and DEC coordinates.
- But there is **precession of the equinox**; i.e., the position of the vernal equinox changes with time



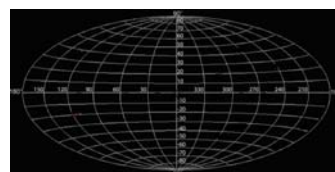
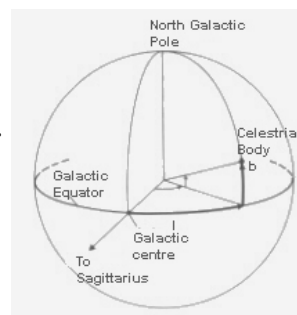
The Earth's spin axis precesses with a period of ~ 26,000 years. By AD13,000 the Polaris will become the star Vega.

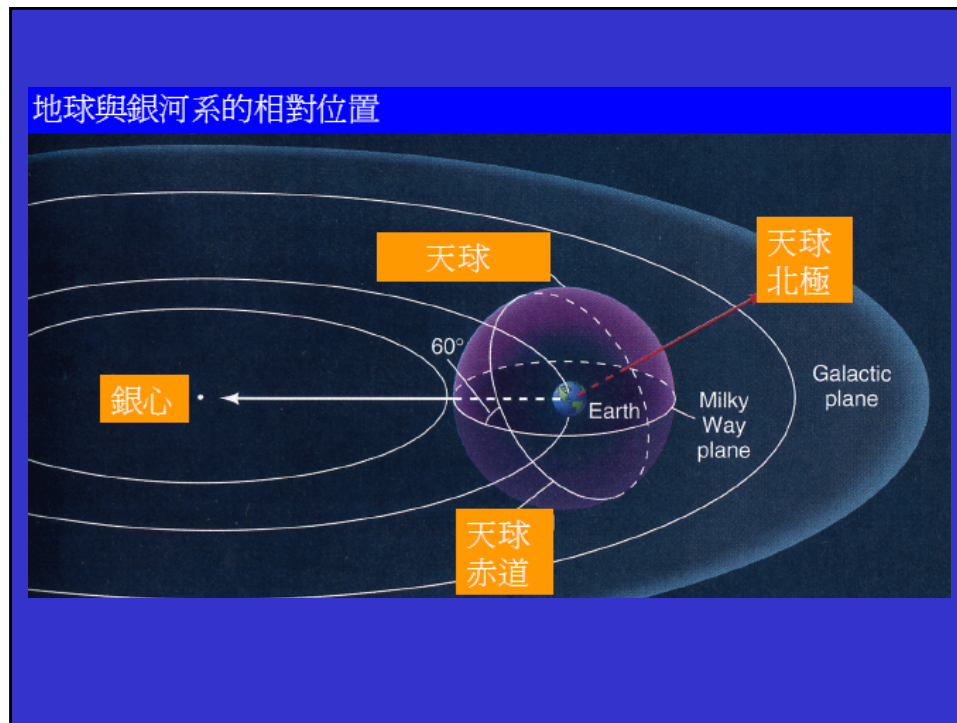


A star's RA and DEC coordinates change with time! B1950 equinox vs J2000 equinox

Coordinate Systems

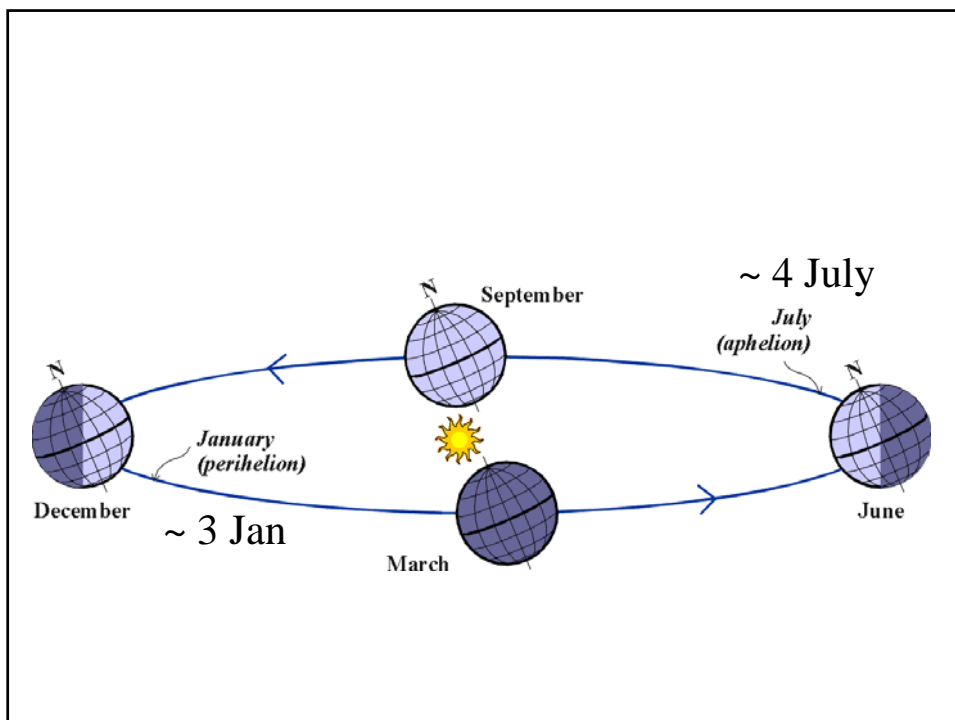
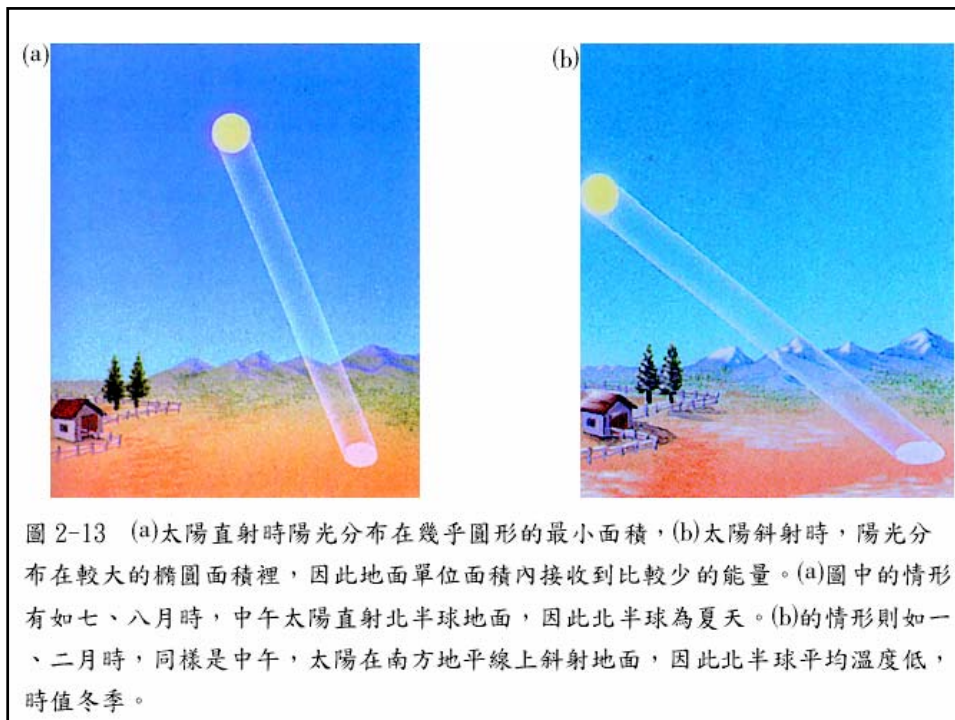
- **Horizontal (observer) system**
- **Equatorial system**
Fundamental plane=earth's equator
- **Galactic system**
Fundamental plane=Galactic plane
- **Ecliptic system**
Fundamental plane=ecliptic plane





Q : What causes the seasoning ?

1. The sunspot activities on the surface of the Sun
2. Inclination of the Earth's spin axis
3. Change of distance between the Sun and the Earth
4. Periodical changes of atmospheric turbulences



- 春分 (vernal equinox)、夏至 (summer solstice)
秋分 (autumn equinox)、冬至 (winter solstice)
- Sun at vernal equinox around March 21 each year
- Direct sunshine → summer; longer days than nights
- Opposite season in southern hemisphere
- At equinoxes, same duration of day and night
- Seasoning has nothing to do with sun-earth distance (difference only 2%).
- In fact Earth at **perihelion** (近日點) in January, at **aphelion** (遠日點) in July

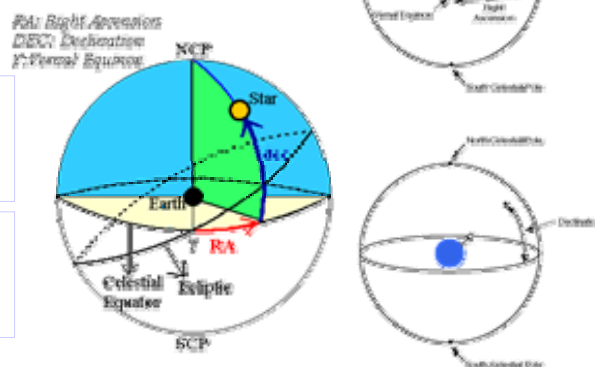
Time and Coordinates

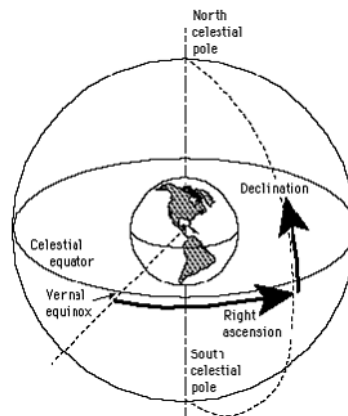
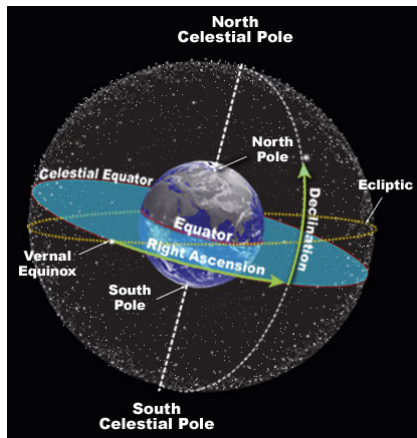
- Sun at vernal equinox on March 21
- Vernal Equinox → RA = 0h
- Sun moves RA ~ 2 hours each month

DEC → 緯度

How does a star of DEC = +25 deg move on the sky?

How about a star of DEC = +90 deg?
DEC = -65 deg?





Remember: Sun at RA=0h around March 21 (vernal equinox 春分), and moves 2 h every month.

An exercise....

- The center of the Milky Way galaxy has the equatorial coordinates of RA~17h and DEC~ -30 deg. What is the best month/season to observe the Galactic center (銀河中心) from Taiwan?
- Is it better to observe the Galactic center in Australia? If so, why? If not, why not?